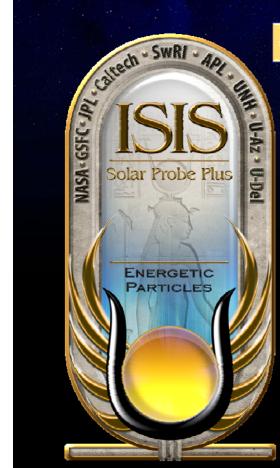
Solar Probe Plus

A NASA Mission to Touch the Sun

Integrated Science Investigation of the Sun Energetic Particles



Preliminary Design Review 05 – 06 NOV 2013

EPI-Hi Mechanical

Sandy Shuman

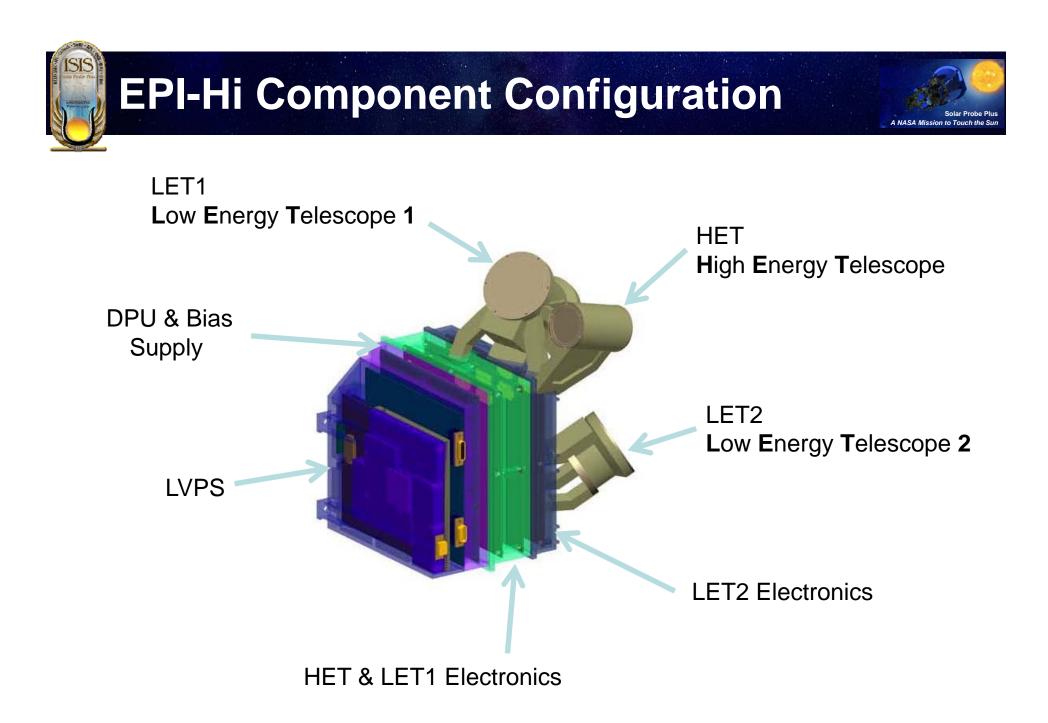
EPI-Hi Mechanical (GSFC)

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EPI-Hi Outline



- Overview of Instrument Configuration
- Location on Spacecraft
- Fields of View
- Mass Allocation
- Mechanical Design
- Assembly Process
- Summary of Peer Review Results



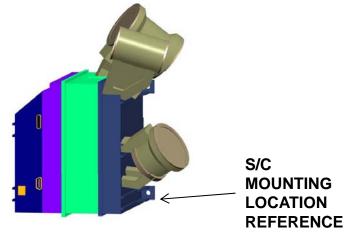
EPI-Hi Instrument

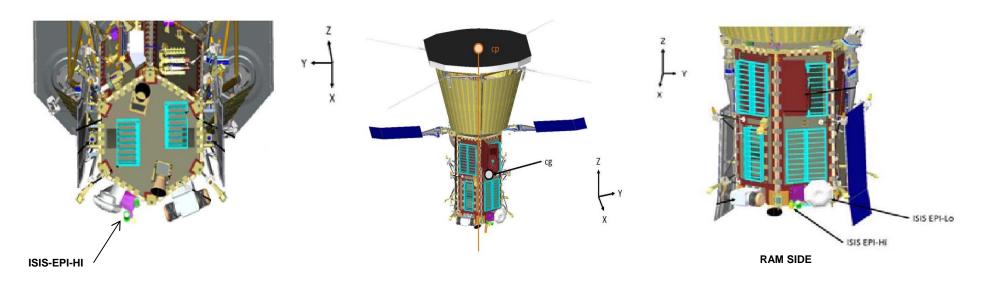


- Location of instrument on spacecraft
 - Located on +X side (RAM side)
 - Lower right mounting bolt on instrument located at: X = 46,16 cm

$$Y = 23,39$$
 cm

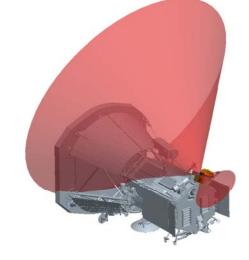
$$Z = 6,13$$
 cm



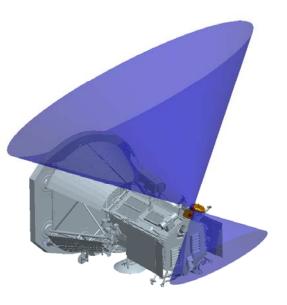


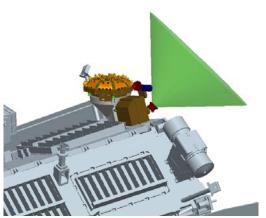
EPI-Hi Instrument FOVs





- HET conical 90° FOV
 - Double-ended
 - 20° above S/C-Sun line
- LET1 conical 90° FOV
 - Double-ended
 - 45° above the S/C-Sun line
- LET2 conical 90° FOV
 - Single-ended
 - Orthogonal to LET1 telescope (135° from S/C-Sun line)







EPI-Hi Mass Allocation



Subsystem	Mass [g]
LET1 telescope	225
LET1 board	258
LET2 telescope	145
LET2 board	233
HET telescope	120
HET board	250
DPU board	197
Bias Supply & RF shields	225 + 130
LVPS & RF shields	160 + 100
Elec. box, hardware & shielding	925 + 250 +100
Telescope brackets	160
Thermal hardware	50
MLI blankets	100
Total	3,628

EPI-Hi Enclosure Requirements



- Work within tight mass constraints
- Design to meet S/C launch environment requirements for Vibration, Acoustics and Thermal conditions
- Design for radiation dose shielding environment
- Package boards maintaining adequate parts clearance boardto-board
- Provide adequate RF and/or ground shielding board-to-board and through the enclosure
- Provide thermal isolation between electronics box and bracket, as well as between telescopes and electronics box

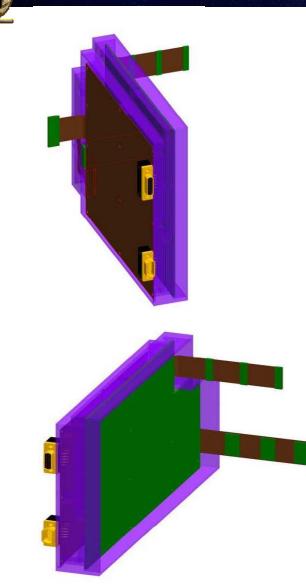
EPI-Hi Electronics Enclosure (1/4)



- Electronics box is made up of 4 major components:
 - LVPS Assembly
 - DPU & Bias Supply Assembly
 - HET & LET1 Electronics Assembly
 - LET2 Electronics Assembly
- Each Electronics Assembly is mounted in a perimeter style frame
- All "frames" when assembled together will provide a continuous RF shield for internal electronics
- Wall thickness will be minimum 1,0 mm (~40 mils) for radiation dose shielding
- Internal shielding between critical components will create separate shielded areas as necessary for proper electronics function
- Board interconnect is achieved using rigid/flex boards with built-in cables terminating to individual nanonics/microstrip connectors on mating boards
- Connections to the S/C will be via standard Micro-D connectors

EPI-Hi Electronics Enclosure (2/4)

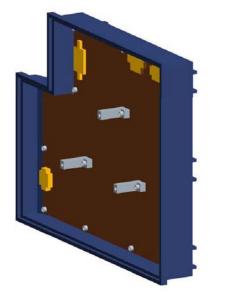




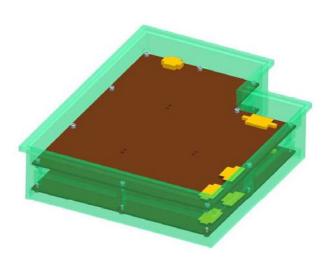
- DPU Board (mounted in one side of frame)
 - Flex connection to telescope boards
 - Flex connection to LVPS
 - S/C cmd&data connector (PCB mount)
 - Thermal harness connector (PCB mt)
 - PCB mounted to machined-in posts in chassis
- Bias Supply Board (mounted in one side of frame)
 - Flex connection to 3 telescope boards
 - Flex connection to DPU board
 - R/F shielding
 - PCB mounted to machined-in posts in chassis

EPI-Hi Electronics Enclosure (3/4)





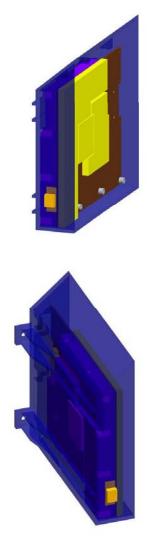
- LET2 Telescope Electronics Assembly
 - Receives flex connection from Bias Supply
 - Receives flex connection from DPU Board
 - Receives 2 flex connections from telescope
 - Housing provides feet for Instrument to enable bracket mounting



- HET & LET1 Electronics Assembly (each board)
 - Flex connection from Bias Supply Board
 - Flex connection from DPU Board
 - I 2 flex connections from its telescope
 - PCB mounted to machined-in posts in chassis

EPI-Hi Electronics Enclosure (4/4)





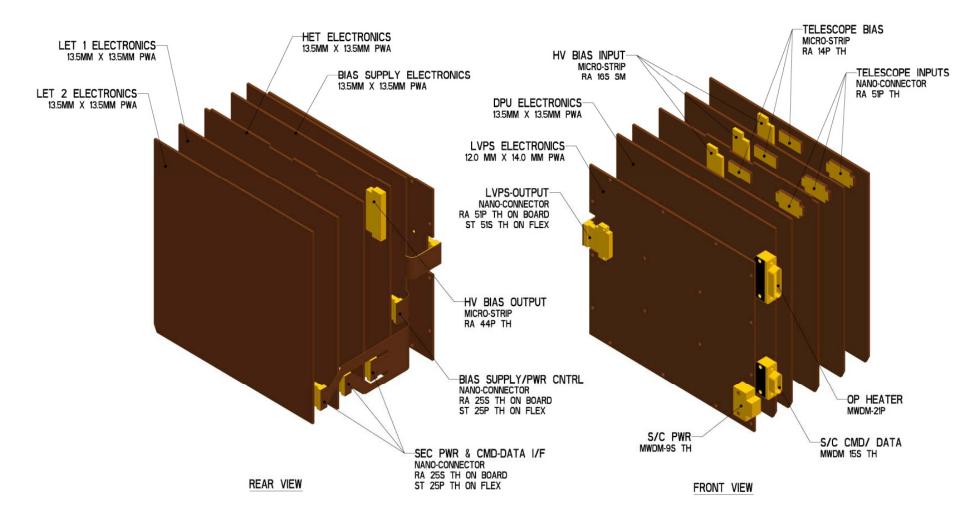
- Low Voltage Power Supply
 - Receives flex connection from DPU Board
 - S/C power connector (PCB mount)
 - Individually shielded primary/secondary circuits top and bottom
 - Housing is tapered to avoid HET FOV
 - Housing provides feet for Instrument to enable bracket mounting
 - PCB mounted to machined-in posts in chassis

*LVPS Board provided by APL *Chassis and shields designed/provided by GSFC



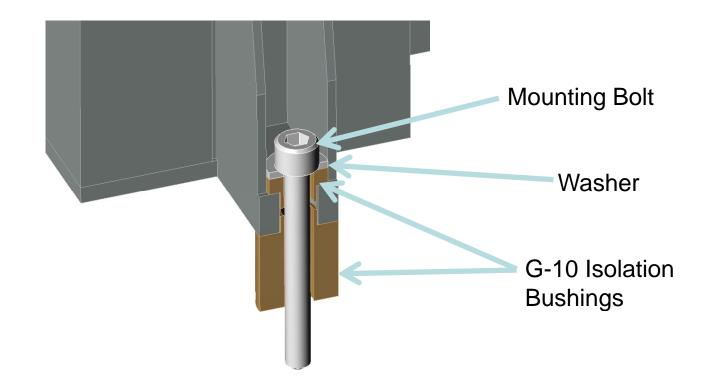
EPI-Hi Interconnect







Typical Mounting Foot Showing Thermal Isolation



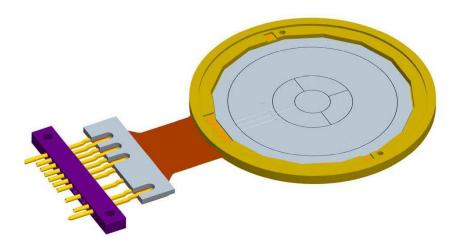
EPI-Hi Mount Design Requirements

Detector Mounts

- Able to transmit signals from silicon detectors, via wire bond connections to output connector
- Allows the stacking of detectors maintaining 0,5 mm spacing surface-to-surface between thickest detectors (1,0 mm)
- Allows any detector to be stacked face-up or face-down with any other detector
- Allows for the protection of wire bonds from being crushed on either side when placed on flat surface during storage and/or test
- Provides electrical breakdown protection next to detector, when stacked, of up to 200 V differential between crown of HV wire bonds to conductive surface of opposing detector

EPI-Hi Detector Mount (1/2)

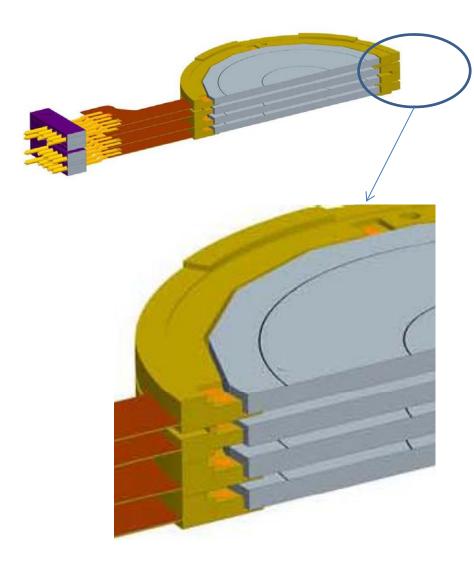




- Recessed detector shelf for silicon detector installation
- Micro-strip connector output
- Flex stiffener to rigidize the area where connector is mounted
- Alignment achieved with alignment pins and concentric stacking shelves on mount and connector
- Tolerancing for mounts will be tightly constrained, but within current CNC machining capabilities
- Detector alignment will be verified through measurement and testing on assembled flight detectors

EPI-Hi Detector Mount (2/2)





- Mount design allows stacking of detectors face-to-face, face-toback and back-to-back while maintaining same spacing
- Mounts are spaced 1,5 mm apart when stacked allowing for 0,5 mm separation between thickest detectors
- Detector voltage ranges from ~2 V up to ~200 V
- Mounts provide adequate spacing/protection for wire bond clearance

EPI-Hi Telescope Design (1/2)

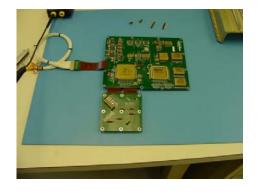


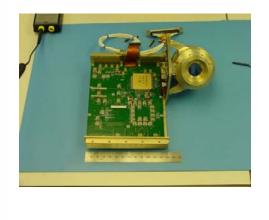
- 3 telescopes comprised of silicon wafer detectors
- Provides ~6,0 mm of aluminum shielding to block unwanted particles from entering through the housing body
- Will have multiple foils for micro-meteorite/light protection
- Mounted directly to the top of the enclosure, allowing the flex interconnect cable to be routed internally to provide proper RF shielding
- Will be thermally isolated from the electronics enclosure
- Will all have red-tag covers over all aperture openings

EPI-Hi Telescope Design (2/2)









- Heritage design
- Uses alignment pins to stack detectors in telescope body
- Mounting bracket designed into telescope body
- Output signal cable will be completely enclosed in assembly, providing proper shielding

Pictures shown are of STEREO\HET telescope

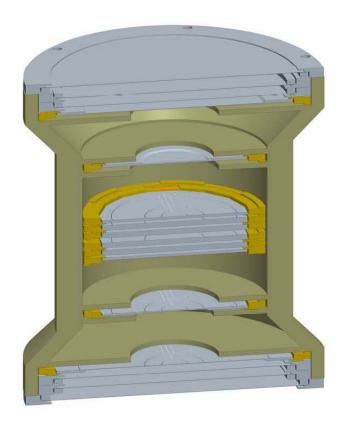
EPI-Hi <u>High Energy Telescope</u>





- 2-~127 µm (5 mil) Foils for micrometeorite/light protection on each end
- Comprised of 16 silicon wafer detectors mounted in rigid-flex mounts
- The front two detectors at each end are spaced apart in order to set a 90° FOV angle

EPI-Hi Low Energy Telescopes



(LET1 shown)

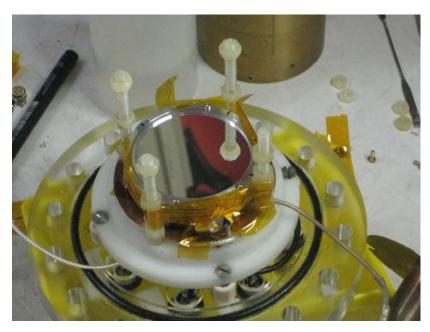
LET1

- Double-ended FOV
- 3 foils for micro-meteorite/light protection on each end
 - Outer foil to be 2 µm polyimide
 - Inner 2 foils to be 1 µm polyimide
- Comprised of 10 silicon wafer detectors mounted in flex-rigid mounts
- The front 3 detectors at each end are spaced apart in order to set a 90° FOV angle
- LET2
 - Single-ended FOV
 - Comprised of one half of an LET1 telescope

EPI-Hi LET Foils (1/2)



- All foils will be aluminized polyimide manufactured by the Luxel corporation
- Full sized prototype foils (1, 2, and 4 micron) have been manufactured by Luxel during Phase B
- Prototype foils have been thoroughly tested, including a high-velocity dust test at the Heidelberg dust accelerator

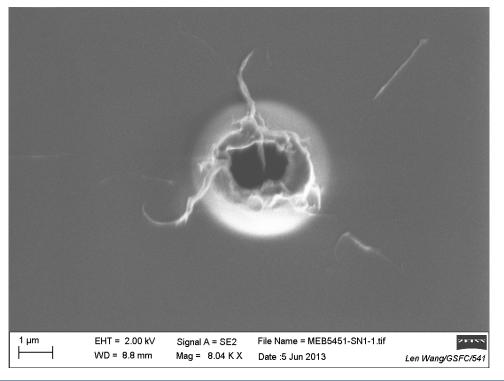


Stack of three Luxel foils (1 micron, 2 micron and 4 micron) in dust accelerator set up

EPI-Hi LET Foils (2/2)



- Dust test shows that holes do not propagate
- Melted polyimide actually appears to strengthen the edge of the hole
- Thermal requirements met with aluminization on the inside surface only

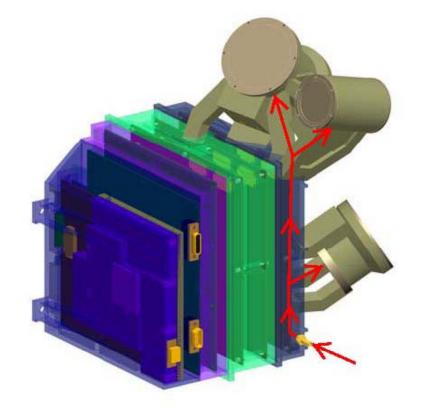


Atomic Force Microscope image from back (exit) side of dust impact in 1 micron thick polyimide



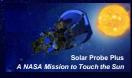
EPI-Hi Telescope Purge

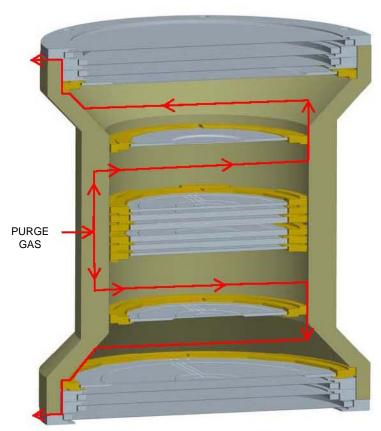




- Purge established to the individual telescopes with a single purge fitting on the outside of the Instrument
- Purge distributed internally through a manifold that sends the purge gas into the center volume of each telescope

EPI-Hi Telescope Venting





- Heritage venting strategy used on several prior missions
- Purge gas enters thru housing into open center volumes
- Gas then flows outwards thru vent slots in housing shelves, detector mounts and foil rings
- Gas exits each end of the telescope thru vent slots below outer foil

EPI-Hi Assembly Flow (1/3)



Electronics Assembly Flow

- Assembled board put into its corresponding frame
- Electronics boards tested independently
- Boards then interconnected and fanned out like a book for troubleshooting and further testing w/ telescopes attached
- Purge hoses and fittings installed
- External RF shields added before assembly is closed up
- Frames then bolted together, and the last remaining board cabling installed/connected through access panels in frames
- Access panel covers installed
- Test, Test, Test

EPI-Hi Assembly Flow (2/3)



Telescope Assembly Flow

- Processed Silicon wafers placed in mounts and tested
- Detector selections made and mount thicknesses recorded
- Detectors stacked in telescope w/ proper shims, covers and spacers
- Polyimide Foils installed in collimators
- Collimators/covers installed onto telescope
- Red Tag/Protective covers installed
- Telescope tested w/ electronics and radioactive sources
- Stored for integration to box

EPI-Hi Assembly Flow (3/3)



Telescope to Electronics Box Assembly

- Mating cable assembled over detector pins and secured in place
- Closeout cover installed over cable
- Telescope positioned over electronics, and cable fed through corresponding frame
- Cables connected at electronics end through access panels in frames
- Access covers installed
- Test, Test, Test
- Telescopes mapped by source testing/accelerator calibrations

EPI-Hi Peer Review Results



- Peer Review conducted earlier this month
- Only 3 issues noted:
 - Thin detectors and implications of environments
 - These have been considered and appropriate testing has been or will be performed
 - PCB/wall-mounted connectors
 - Appropriate measures will be taken to minimize stresses during installation
 - Whether "bolt slip" during instrument/telescope mounting will be sufficient enough to keep telescope FOVs within spec
 - This will be analyzed and verified



Summary



- Mechanical concept verified with peer review
- All issues from peer review addressed
- Next step, the drawings for Engineering Model, already started